

Entropy Change of Black Holes of Quarter Spin in XRBs & AGN

¹Rakesh Paswan, ²Niharika Kumari, ³Varsha Kumari and ⁴Dipo Mahto

Author's Affiliations:	¹ Assistant Professor & HOD, Dept. of Mathematics, DSM College Jhajha, India ² Lecturer, Department of Physics, Govt. Polytechnic Khagaria, India ³ Department of Physics, Indian Institute of Engineering, Science and Technology, Sibpur, Kolkata, India ⁴ Professor & Head, Department of Physics, Bhagalpur College of Engineering, India
*Corresponding author:	Dipo Mahto Professor & Head, Dept. of Physics, Bhagalpur College of Engineering, Bhagalpur, India E-mail: dipomahto@hotmail.com
ABSTRACT	This paper is concerned with the study of change in entropy of black holes of spin parameters $a^* = +1/4$ & $-1/4$ for unit angular velocity (Ω) for XRBs and AGN. This result that the black holes of co-rotation as well as counter rotation decreases and increases the entropy respectively and the emission of energy from BHs is predicted due to loss of mass as Hawking radiation on the basis of quantum mechanical fields.
KEYWORDS	Quarter Spin, XRBs and AGN.

Received on 27.04.2025, Revised on 14.06.2025, Accepted on 10.07.2025

How to cite this article: Paswan R., Kumari N., Kumari V. and Mahto D. Entropy Change of Black Holes of Quarter Spin in XRBs & AGN. *Bulletin of Pure and Applied Sciences- Physics*, 2025;44D (2): 56-60.

INTRODUCTION

The classical theory is not able to explain the emission of radiation from BHs. The present theory explains only capturing of radiation or any information near the event horizon. The quantum theory gives the explanation of emission of radiation a black body radiation. In this case the temperature and surface gravity are proportionality related [1]. Many works similar to this work have been done [2, 3, 4, 5, 6, 7, 8, 9, 10, 11, 12, 13]. Rakesh et al. discussed the change in entropy of Bosonic Field of Black Holes of

spin $a^* = +3$ & -3 in XRBs and AGN [14], whereas, Mahto et al also discussed the same for spin parameters $a^* = +2$ & -2 [15],

The work deals a model about the change in entropy of spinning black holes w.r.t. the mass of spin parameters $a^* = +1/4$ & $-1/4$ for unit angular velocity (Ω) in X R Bs and A G N.

THEORETICAL DISCUSSION

The entropy change with respect to the mass is given by the following equation [8].

$$\delta S / \delta M = 8\pi M(1 - 2\Omega M a^* + a^{*2} / 2 - M\Omega a^{*3}) \quad (1)$$

Let us apply this model for $a^*=+1/4$ and $a^*=-1/4$

[16] and unit angular velocity ($\Omega =1$), the equation (1) becomes

$$\left(\frac{\delta S}{\delta M}\right)_{+1/4} = -\frac{33\pi M}{8}(M - 2) \quad (2)$$

$$\left(\frac{\delta S}{\delta M}\right)_{-1/4} = \frac{33\pi M}{8}(M + 2) \quad (3)$$

The condition for max./min. is given as.

$$\left(\frac{\delta S}{\delta M}\right)_{\pm 1/4} = 0 \quad (4)$$

Using the above equations on (2) and (3), we have

$$M(M - 2) = 0 \quad (5)$$

$$M(M + 2) = 0 \quad (6)$$

When the equations(5) and (6) are solved to give

$$M=0 \text{ or } M=2 \quad (7)$$

$$M=0 \text{ or } M=-2 \quad (8)$$

Now the models represented by the equations (2) and (3) are applied for the different masses of black holes in XRBs and AGN and also tabulated in the tables

Table 1: Entropy change of BHs of spin +1/4 and -1/4 in XRBs.

S. N.	Mass in (M_{\odot})	$\left(\frac{\delta S}{\delta M}\right)_{+1/4} = -\frac{33\pi M}{8}(M - 2)$ [Joule/Kelvin/kg] $\times 10^{63}$	$\left(\frac{\delta S}{\delta M}\right)_{-1/4} = \frac{33\pi M}{8}(M + 2)$ [Joule/Kelvin/kg] $\times 10^{63}$
1	5 M_{\odot}	-1 . 29525	+1 . 29525
2	6 M_{\odot}	-1 . 86516	+ 1 . 86516
3	7 M_{\odot}	-2 . 53859	+ 2 . 53859
4	8 M_{\odot}	- 3 . 31584	+ 3 . 31584
5	9 M_{\odot}	- 4 . 19661	+ 4 . 19661
6	10 M_{\odot}	- 5 . 18100	+ 5 . 18100
7	11 M_{\odot}	- 6 . 26901	+ 6 . 26901
8	12 M_{\odot}	- 7 . 46064	+ 7 . 46064
9	13 M_{\odot}	- 8 . 75589	+ 8 . 75589
10	14 M_{\odot}	- 10 . 1548	+ 10 . 1548
11	15 M_{\odot}	- 11 . 6573	+ 11 . 6573
12	16 M_{\odot}	- 13 . 2634	+ 13 . 2634
13	17 M_{\odot}	- 14 . 9731	+ 14 . 9731
14	18 M_{\odot}	- 16 . 7864	+ 16 . 7864
15	19 M_{\odot}	- 18 . 7034	+ 18 . 7034
16	20 M_{\odot}	- 20 . 7240	+ 20 . 7240

Table 2: The change in entropy of BHs of spin +1/4 and -1/4 in AGN.

S. N.	Mass in M_{\odot}	Mass in $10^7 M_{\odot}$	$\left(\frac{\delta S}{\delta M}\right)_{+1/4} = -\frac{33\pi M}{8}(M-2)$ [Joule/Kelvin/kg] $\times 10^{77}$	$\left(\frac{\delta S}{\delta M}\right)_{-1/4} = \frac{33\pi M}{8}(M+2)$ [Joule/Kelvin/kg] $\times 10^{77}$
1	$1 \times 10^6 M_{\odot}$	0.1	-0.0005181	+ 0.0005181
2	$2 \times 10^6 M_{\odot}$	0.2	-0.0020724	+ 0.0020724
3	$3 \times 10^6 M_{\odot}$	0.3	-0.0046629	+ 0.0046629
4	$4 \times 10^6 M_{\odot}$	0.4	-0.0082896	+ 0.0082896
5	$5 \times 10^6 M_{\odot}$	0.5	-0.0129525	+ 0.0129525
6	$6 \times 10^6 M_{\odot}$	0.6	-0.0186516	+ 0.0186516
7	$7 \times 10^6 M_{\odot}$	0.7	-0.0253869	+ 0.0253869
8	$8 \times 10^6 M_{\odot}$	0.8	-0.0331584	+ 0.0331584
9	$9 \times 10^6 M_{\odot}$	0.9	-0.0419661	+ 0.0419661
10	$1 \times 10^7 M_{\odot}$	1.0	-0.0518100	+ 0.0518100
11	$2 \times 10^7 M_{\odot}$	2.0	-0.2072400	+ 0.2072400
12	$3 \times 10^7 M_{\odot}$	3.0	-0.4662900	+ 0.4662900
13	$4 \times 10^7 M_{\odot}$	4.0	-0.8289600	+ 0.8289600
14	$5 \times 10^7 M_{\odot}$	5.0	-1.2952500	+ 1.2952500
15	$6 \times 10^7 M_{\odot}$	6.0	-1.8651600	+ 1.8651600
16	$7 \times 10^7 M_{\odot}$	7.0	-2.5386900	+ 2.5386900
17	$8 \times 10^7 M_{\odot}$	8.0	-3.3158400	+ 3.3158400
18	$9 \times 10^7 M_{\odot}$	9.0	-4.1966100	+ 4.1966100
19	$1 \times 10^8 M_{\odot}$	10.0	-5.1810000	+ 5.1810000
20	$2 \times 10^8 M_{\odot}$	20.0	-20.7240000	+ 20.7240000
21	$3 \times 10^8 M_{\odot}$	30.0	-46.6290000	+ 46.6290000
22	$4 \times 10^8 M_{\odot}$	40.0	-82.8960000	+ 82.8960000
23	$5 \times 10^8 M_{\odot}$	50.0	-129.5250000	+ 129.5250000
24	$6 \times 10^8 M_{\odot}$	60.0	-186.5160000	+ 186.5160000
25	$7 \times 10^8 M_{\odot}$	70.0	-253.8690000	+ 253.8690000
26	$8 \times 10^8 M_{\odot}$	80.0	-331.5840000	+ 331.5840000
27	$9 \times 10^8 M_{\odot}$	90.0	-419.6610000	+ 419.6610000
28	$1 \times 10^9 M_{\odot}$	100.0	-518.1000000	+ 518.1000000

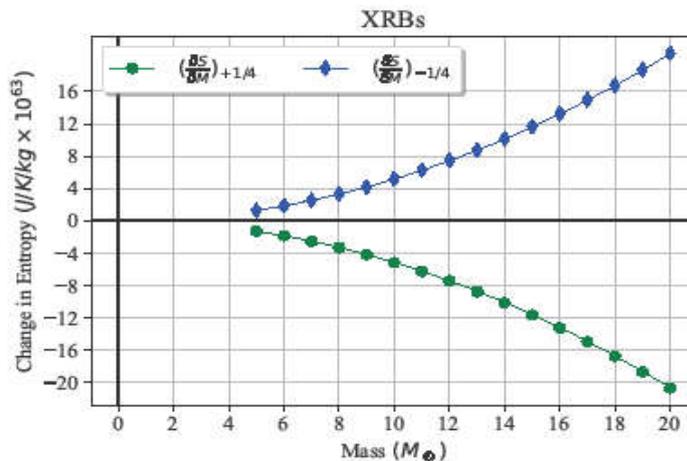


Figure 1: Entropy change of BHs of spin parameter +1/4 and -1/4 in XRBs.

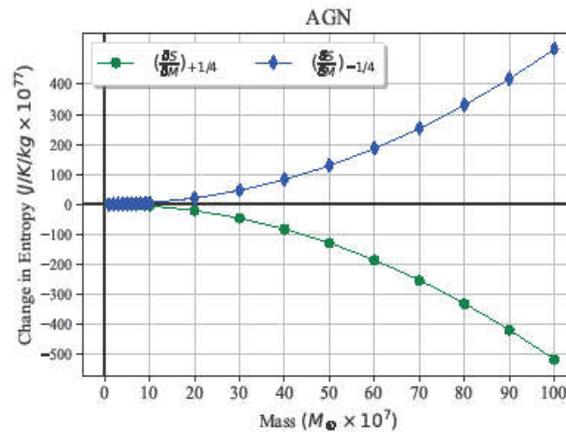


Figure 2: Change in entropy of BHs of spin +1/4 and -1/4 in AGN.

RESULT & DISCUSSIONS

The work starts to consider a model represented by the equation

$$\delta S / \delta M = 8\pi M (1 - 2\Omega M a^* + a^{*2} / 2 - M \Omega a^{*3})$$

. The model is applied for spin $a^* = +1/4$ and $-1/4$ to give two specific models represented by the equations (2) & (3). These two equations will provide us two specific relations for quarter spin of black holes. To get the maximum/minimum values, the equations (2) and (3) are conditioned to the equation (4) to give equations (5) &(6) respectively and then solved to get $M = 0, +2$ and -2 . Now these models are used for the test masses of BHs in XRBs and AGN and also tabulated in the tables 1 & 2. The graphs are also plotted for the different masses of black holes using equations (2) and (3) represented by figures 1for XRBs and 2 for AGN.

The observations from the tables and graphs show that the variation for change in entropy w.r.t. mass are the same and gets symmetrical form for both rotations for each category.

From the observations of our data calculated in the present work, we also see that that there are three types of masses in which one is zero and other two values are positive (+) and negative (-). The significances of these signs (+)and (-) have the basic importance. The zero mass of

photon is justified from our work. The sign (+) gives the classical concept regarding the mass. The negative sign (-) explains naked singularity and also dark matter.

CONCLUSIONS

The black holes of quarter spin for both rotations the prediction the emission of Hawking radiation due to reduction of mass on the basis of quantum mechanical fields along with naked singularity of black and dark matter.

REFERENCES

1. Hawking SW. Particle creation by black holes. **Commun Math Phys.** 1975;43:199–220.
2. Bardeen JM, Carter B, Hawking SW. The four laws of black hole mechanics. **Commun Math Phys.** 1973;31(2):161–170.
3. Bekenstein JD. Bekenstein–Hawking entropy. **Scholarpedia.** 2008;3:7375.
4. Carlip S. Black hole thermodynamics and statistical mechanics. In: **Les Houches Summer School Proceedings.** 2009;769:89–123.
5. Dabholkar A. Black hole entropy in string theory: A window into the quantum structure of gravity. **Curr Sci.** 2005;89(12):25.
6. Mahto D, Kumari K, Sah RK, Singh KM. Study of non-spinning black holes with

- reference to the change in energy and entropy. **Astrophys Space Sci.** 2012;337:685-691.
7. Mahto D, Singh AK, Ram M, Vineeta K. Change in internal energy and enthalpy of the black holes. **Int J Astrophys Space Sci.** 2013;1(4).
 8. Mahto D, Kumari A. Change in entropy of spinning black holes due to corresponding change in mass in XRBs. **Int J Astron Astrophys.** 2018;8:171-177.
 9. Mahto D, Paswan R, Kumari K, Kumar B. Change in entropy of fermionic fields of black holes with respect to mass. **J Inf Comput Sci.** 2020;10(9):342-351.
 10. Mahto D, Paswan R, Kumar B, Kumari A. Change in entropy of bosonic fields of black holes with respect to the mass for maximum co-rotation and counter-rotation. **J Inf Comput Sci.** 2020;10(9):357-369.
 11. Narayan R. Black holes in astrophysics. **New J Phys.** 2005;7(1):1-31. Available from: arXiv:gr-qc/0506078
 12. Strominger A, Vafa C. Microscopic origin of Bekenstein-Hawking entropy. **Phys Lett B.** 1996;379:99-104.
 13. Traschen J. An introduction to black hole evaporation [Internet]. arXiv preprint; 2000. Available from: arXiv:gr-qc/0010055
 14. Paswan R, Kumari N, Kumar B, Mahto D. A study on mathematical model of entropy change of bosonic field of black holes of spin parameters ($a^* = +3, -3$) in XRBs and AGN. **J Harbin Eng Univ.** 2024;45(12):127-132.
 15. Mahto D, Paswan R, Kumari N, Kumar B. A study of mathematical model on entropy change of bosonic field of black holes of spin parameters $a^* = +2$ & -2 in XRBs and AGN. **J Harbin Eng Univ.** 2024;45(12):133-138.
 16. Yash. An introduction to spin in quantum mechanics (Spin: Explained). **Quantaphy.** 2022.
